Segregation scheme of indium in AlGaInAs nanowire shells

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Quaternary alloys enable the independent optimization of different semiconductor properties, such as the separate tuning of the band gap and the lattice constant. Nanowire core-shell structures should allow a larger range of compositional tuning as strain can be accommodated in a more effective manner than in thin films. Still, the faceted structure of the nanowire may lead to local segregation effects. Here, we explore the incorporation of indium in AlGaAs shells up to 25%. In particular, we show the effect of In incorporation on the energy shift of the AlGaInAs single-photon emitters present in the shell. We observe a redshift up to 300 meV as a function of the group-III site fraction of In. We correlate the shift with segregation at the nanoscale. We find evidence of the segregation of the group-III elements at different positions in the nanowire, not observed before. We propose a model that takes into account the strain distribution in the nanowire shell and the adatom diffusion on the nanowire facets to explain the observations. This work provides novel insights on the segregation phenomena necessary to engineer the composition of multidinary alloys.

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Introduction. Quaternary alloys enable to independently tune the semiconductor lattice constant and band gap by careful composition engineering. However, the controlled deposition of quaternary alloys with randomly distributed composition can be challenging. Miscibility gaps and diffusion may induce the segregation of the different species that compose the alloy, and limit the possible compositions [1,2]. Moreover, strain may build up when quaternary semiconductors epitaxially grow on substrates of different lattice constant. To avoid plastic relaxation [3,4], the theoretical range of available compositions is significantly narrowed in real applications. Due to their reduced diameter, semiconductor nanowires (NWs) provide a suitable growth platform to minimize plastic relaxation, [5-10] permitting lattice-mismatched material combinations not achievable by planar schemes [11-14].

In this work, we explore the range in which In can be incorporated in AlGaAs shells of core-shell GaAs-AlGaAs NWs. The goal is to understand how much the emission energy of single-photon emitters, spontaneously forming in the shell, can be redshifted [15-17]. In particular, we want to dissociate the band-gap engineering from additional segregation effects that may occur due to the presence of In. Consequently, we characterize the optical emission and correlate it with the incorporation of In using several techniques at different length scales. We find that the group-III elements in the shell segregate to different positions and generate regions of different band gaps. In particular, we find consistent evidence of the In segregation in novel wedge-shaped In-rich features. We correlate these observations with the threefold polarity of the NW cross section and the migration of the adatoms on the NW sidewalls. In addition to the role of the crystal directions of the NW facets [18-20], we consider strain to cause the peculiar In segregation; the presence and distribution of strain are further analyzed by finite-element simulations as well as Raman and photoluminescence (PL) spectroscopy.

The material system: core-shell NWs. The GaAs-AlGa(1−x)In(x)As core-shell NWs are grown in a high-mobility molecular-beam-epitaxy system (MBE, DCA P600). We use Si(111) substrates and self-catalyzed growth [23,24] at a substrate temperature of 640 °C. When the NWs are about 10-μm long, we stop the axial growth by interrupting the Ga supply and consuming the catalyst. We then lower the substrate temperature to 460 °C to promote radial growth of an AlGa(In)As shell on the NW sidewalls [25,26]. We start with an AlxGa1-xAs shell and then incorporate In without modifying the Ga or Al rates. For this, we varied the In partial pressure from 4.5 × 10−9 to 2.5 × 10−7 Torr, while the Al and Ga pressures were kept constant. Unless stated, the samples are identified by the group-III site fraction of In measured through various techniques and expressed as percentage in the notation AlxGa1−xInyAs (Al, Ga, and In sum to 100%): 0%, 1%, 2%, 3%, 4%, 15%, and 25%. An outer GaAs shell of 5 nm prevents the oxidation of the inner AlGa(In)As.

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Previous reports show that the composition of an AlGaAs shell deposited around a GaAs NW exhibits nanoscale fluctuations [15,16,27–30]. This phenomenon depends on the different surface mobility of Ga and Al during growth and the presence of surface-potential gradients \((\mu)\) on the NW sidewalls. The formation of Al-rich planes at the ridge between two NW facets was reported [19]. Furthermore, Ga-rich clusters can form quantum-dots (QDs) that behave as single-photon emitters in the NW shell [15,16,29]. A sketch of a core-shell GaAs-AlGaAs NW and its cross section is presented in Fig. 1(a), illustrating the distribution of Al-rich planes and QDs made from Ga-rich nanoclusters.

**Light-emission properties.** First we address the optical functionality of the core-shell NWs and embedded QDs. We used a single-frequency optically pumped semiconductor laser, with a wavelength of 532 nm, focused in a spot of less than 1 \(\mu\)m in diameter (100 W/cm²) to measure NW samples at 12 K using a helium cryostat. The PL signal is collected into a spectrometer and dispersed by a 300 l/mm grating onto a Peltier-cooled CCD. The QDs in the AlGaAs shell emit bright and narrow PL lines (linewidth below 100 \(\mu\)eV) [15] between 1.7 and 1.9 eV [27]. The green spectrum at the bottom of Fig. 1(b) illustrates the PL emission of these structures; the sharp peaks are attributed to the presence of QD single-photon emitters [15,31]. Figure 1(b) also contains PL spectra of the shell-QD emission for increasing In fraction, revealing several sharp peaks at different energies. The emission-energy range redshifts with increasing In fraction in the shell. One can also qualitatively determine that the QD emission linewidth broadens in samples of higher In fraction. The sample with the highest In fraction showed a decreased PL intensity, compensated for by increasing the laser power [see Fig. 1(b)].

For a statistical analysis on large ensembles of NWs and QDs, Fig. 1(c) shows the median emission energy of the QDs as a function of the measured In pressure in the MBE chamber (bottom axis) or as a function of the In fraction (top axis) measured in the shell by x-ray fluorescence (XRF) combined with atom probe tomography (APT) for 1% to 4% of In and scanning transmission electron microscopy energy-dispersive x-ray spectroscopy (STEM EDX) for 15% and 25% In, see Ref. [32]. We measured 25 NWs from each sample. The QD emission lines were identified by an automatic routine [27]. The dashed-dotted line indicates the expected shell band gap as a function of shell composition [21]. The dashed line corresponds to the GaAs band gap. Both lines are corrected for the simulated strain [22] that arises from the core-shell lattice mismatch, as further discussed in the manuscript.

The inset in Fig. 1(c) shows a Hanbury-Brown-Twiss autocorrelation measurement \(g^2(\tau)\) of the exciton line of a QD from a sample with 3% In. The power-dependent PL showing the exciton nature of the emission line is reported in the supporting information. The sample is measured at 4.2 K in an Attody 700 closed-cycle cryostat and is excited in continuous wave by a HeNe laser at 632.8 nm through an objective with NA = 0.81, which also collects the QD signal from the cryostat. A 1200 l/mm grating is used to select the QD line of interest that is sent to a 50:50 beam splitter. The two paths of the beam splitter are coupled to single-mode optical fibers that send the signal to two single-photon avalanche diodes, one for each path. The dip in the autocorrelation function in the inset of Fig. 1(c) is below 0.5, which is the signature of the single-photon emission. Several factors increase the count at zero delay, including background counts from QD lines spectrally close to the chosen one.

The effect of the shell-composition engineering on the QD emission energy is significant. From Fig. 1(c), the median QD
emission redshift is visible as a function of the increasing In incorporation, by up to about 300 meV. In a random-distribution alloy, the addition of In to AlGaAs lowers the shell band gap and the QD emission energy. Since lower-band gap nanoscale clusters in the AlGa(In)As matrix form the studied QDs, their emission is always expected to be redshifted with respect to the band gap of the host shell matrix [dashed-dotted line in Fig. 1(e)]. In Fig. 1(c), we observe that the energy difference between the QD emission and the shell band gap is gradually reduced for higher In fractions. This calls for a different segregation effect at low and high In fractions, whose mechanisms and details may be investigated in a further study.

**Compositional analysis at the nanoscale.** In NWs the composition can vary significantly at the nanometer scale due to their faceted nature. The hexagonal cross-section of a GaAs NW exhibits six {110} facets. The ridges between two of these facets are {112} nanofacets with alternating A- and B-polar (Ga- and As-terminated) surfaces [18,19,30]. {110} and {112} facets have different and species-dependent sticking coefficients [30]. In addition, the increased surface curvature at the ridges increases the local surface energy and affects the adatom mobility. Here, we provide compositional measurements with spatial resolution down to the nanometer scale: APT and STEM-based electron energy loss spectroscopy (EELS) and EDX.

We performed laser-assisted APT measurements on the sample with 2% In. The specimen is cooled to 80 K and irradiated with UV laser light (343-nm wavelength) in 2-nJ pulses; the detection rate is 0.0025 events/pulse. The evaporated NW volume is a cylinder with diameter of 64 nm and length of 90 nm; in Figs. 2(a)–2(e), it is shown as a 2D projection on a plane perpendicular to the NW growth axis. Figures 2(a)–2(e) show, respectively, the In, Al, and Ga fractions. It is possible to distinguish the GaAs core by the absence of In and Al. In the shell, radial segregation of Al along the ridges between two facets of the hexagonal NW core is visible as the three Al-rich stripes in the reconstructed NW volume, in agreement with previous works [15,18,20]. In addition to the Al-rich planes, APT reveals that the shell distribution of the other group-III atoms is not perfectly random throughout the NW shell. A slightly higher In fraction (up to 3%) is visible in proximity of two of the Al-rich planes indicated as 1 and 3 in Fig. 2(a). Farther from these positions, the average In fraction decreases to about 1.6%. In Figs. 2(b) and 2(e), we can also observe that the shell in proximity of the Al-rich planes is slightly depleted in Al and enriched in Ga. At shorter length scales, APT-based distributions [16] of the distances between first-neighbor atoms in the shell do not evidence short-range reordering of the quaternary alloy (Fig. S13 in Ref. [32]).

To study the segregation of Al, Ga, and In in more extended regions, we studied the NW cross-sections by high angular annular dark field (HAADF) STEM and EELS in an aberration-corrected (AC) TEM microscope (FEI Titan) operated at 300 keV. We used the sample with 15% In. We prepared NW cross sections to directly map the shell composition: the NWs were spread on a Si substrate and the cross sections were FIB-cut perpendicularly to the NW growth direction ([111]). The cross-section HAADF micrographs are acquired along the (111) zone axis, with the NW sidewalls corresponding to the [110] family (Fig. S4 in Ref. [32]). Figure 2(d) reports the HAADF micrograph of one NW cross section. The HAADF contrast depends on the sample thickness and composition: the higher the atomic number (Z) of the species in the sample, the brighter the HAADF signal. Through high-precision FIB cut, the NW cross sections have negligible thickness variations; the contrast in Fig. 2(d) depends on the average Z at different positions across the sample. In this figure, the hexagonal shape of the NW cross section is clearly visible and the core is distinguishable from the shell. The dark stripes that form in the shell along the ridges between two NW facets ([112] crystalline directions) are due to local enrichment with a light element, such as Al, as already observed by APT and in agreement with the literature [15,18,20]. It is possible to observe that the thickness of the Al-rich planes follows a threefold symmetry as observed by Zheng et al. [18]. We obtain a thickness of 3 and 1.7 nm for the thick and thin planes, respectively (more details in Fig. S4 in Ref. [32]). Similarly, the HAADF contrast shows that the thick planes are richer in Al than the thin ones. Comparing our observations with the literature [18], we assign the thin planes to the A polarity and the thick ones to the B polarity. Accordingly, we labeled the two polarities in Fig. 2(d) as A (orange) and B (blue). We highlight that, without the alternation of the A and B polarity of the {112} nanofacets, all the corners of the hexagonal NW cross-section would be equivalent.

With the exception of the Al-rich planes, the HAADF micrograph in Fig. 2(d) shows little compositional contrast in the shell. By carefully inspecting Fig. 2(d) and Fig. S4 in Ref. [32], a slightly brighter contrast is visible in proximity of the Al-rich planes: it corresponds to a local increase in the fraction of elements with higher Z, such as Ga and In. To clarify this observation, in Figs. 2(e)–2(h), we report the EELS maps of the upper half of the NW cross-section shown in Fig. 2(d): the selected region, indicated by the dashed rectangle in Fig. 2(d), includes an A-polar and two B-polar ridges as well as part of the [110] NW sidewalls. The Ga, In, Al, and As distributions are shown in Figs. 2(e)–2(h), respectively. Principal component analysis (PCA) was used to enhance the signal-to-noise ratio in these maps [33,34]. The NW core is clearly distinguished from the shell by the absence of Al and In and a thin Ga-rich layer, corresponding to the GaAs capping, surrounds the shell [Fig. 2(e)]. As is randomly distributed throughout [Fig. 2(f)], as opposed to the distributions of the group-III elements: the presence of Al-rich planes along the hexagon ridges is confirmed [Fig. 2(g)], together with the Ga and In depletion at the same positions [Figs. 2(e) and 2(f)]. Reports on GaAs-InGaAs and GaAs-InAlAs core-shell NWs [35] also show Ga and In depletion along the [112] direction. The brightness and thickness of the Al-rich planes agree with the threefold polarity-driven segregation previously discussed.

In agreement with the localized In enrichment shown by APT [Fig. 2(a)], STEM EELS shows in Fig. 2(f) In-rich features surrounding the Al-rich planes at the [112]B directions: interestingly, moving from the core to the outer shell, the In-rich regions become broader with a wedge-like shape on each side of the Al-rich B-polar planes. Ga has a more random distribution in the shell, although a slight increase is visible around the A-polar [112] nanofacet in Fig. 2(e). This Ga-rich
feature is not as pronounced as the In-rich ones surrounding the B-polar planes; it also has a relatively constant thickness around the A-polar planes, while the In-rich features develop an unusual wedge-shaped profile. One may also observe that, in proximity of the Al-rich planes aligned in the \(\langle 112\rangle\) directions of both polarities, the Al fraction slightly decreases [Fig. 2(g)]. Although the contrast is not sharp, this would agree with the local increase in the Ga and In fractions at the same positions. EELS maps on whole NW cross sections confirm the threefold symmetry of the shell segregation and the In enrichment only around the three \(\langle 112\rangle\)B directions (see the supporting information). On the left of Fig. 2(i), a scheme summarizes the main compositional segregations observed in the NW cross sections [the scheme is aligned with Figs. 2(d)–2(h)]: the red-shaded areas indicate the wedge-shaped In-rich segregation and blue stripes indicate the alternatively thicker and thinner Al-rich planes. The polarity of the \{112\} nanofacets is labeled.

We acquired STEM-EDX maps of NW cross sections from the same samples as the one used for EELS. We use a FEI Tecnai Osiris TEM operated in STEM mode at 200 kV with a probe current exceeding 1 nA. The x-ray signal is collected by four silicon drift detectors under a solid angle of 0.9 srad. The NW cross sections are prepared by embedding the as-grown sample into epoxy. After hardening, the epoxy with embedded NWs is detached from the growth substrate and is mounted into an ultramicrotome and cut into 80-nm-thick slices transferred on a TEM grid. The EDX maps (Fig. S3 in Ref. [32]) for the Ga, Al, and In distributions in the core-shell cross sections confirm the presence of the same features observed in the EELS maps. Our observations are particularly robust: three independent techniques (APT, EELS, and EDX) confirm the formation of novel threefold wedge-shaped In-rich segregates in NWs with different average In fractions; together with the clear symmetry of these features, the agreement of the three techniques excludes artifacts due to sample preparation.
We now turn to a quantitative analysis of the distribution of the different species in the NW cross section: the EDX-based quantification is more accurate than the one based on EELS [36]. Figure 2(i) shows on the right a plot of the Ga (green curve), Al (blue curve), and In (red curve) atomic percentage extracted from a STEM EDX map as a function of the position along the dashed circumference shown in the sketch on the left (i.e., a circular linescan of 240 nm in length). The data points at 0 and 240 nm correspond to the same position on the circumference (arrow in the sketch). The shaded areas in the plot are colored in orange and blue with the same coding as in Fig. 2(d) to distinguish the A- and B-polarity of the \{112\} facets, respectively. The data points acquired in the middle of the flat \{110\} facets have a white background.

We focus on the most prominent fluctuation around the \langle 112 \rangle B facet at position 100 nm. Moving from the \{110\} plane to the \{112\}B nanofacet, the Al fraction first decreases (from 25% to 20%), then rises to 60% at the corner. In smaller proportions, the In distribution is opposed to the Al trend: for the same positions, In first increases (from 14% to 19%) and then decreases to 12% at the corner. Ga is almost constant at 61% and 60% and then decreases to 28% at the corner. Very similar trends in Al/Ga/In fractions are reproduced at the six corners of the hexagonal NW cross section and are modulated according to a threefold symmetry: the compositional variations are consistently more pronounced in proximity of the B-polar facets.

**Strain analysis and growth model.** We turn now to the understanding of the element distribution in the shell by taking strain into account. Figure 3(a) shows the PL spectra of the GaAs core at 12 K for core-shell NWs of increasing In fraction. The band-edge GaAs PL redshifts from 1.51 eV down to 1.37 and 1.29 eV for shells containing 15% and 25% In, respectively, which we attribute to the tensile strain imposed by the shell. Tensile strain in the core is also confirmed by Raman spectroscopy [32].

To gain insight in the strain distribution imposed by the lattice-mismatched core-shell NW, we simulated the strain given the lattice mismatch, dimensions and geometry of the structure. We use the methodology of Boxberg et al. [37], adapted to hexagonal GaAs-AlGaInAs NWs at low temperature [21] in the software NEXTNANO [38,39]. The strain magnitude and distributions in the core and shell are calculated by minimizing the elastic energy due to the lattice mismatch between the two. Figure 3(b) shows a map of the hydrostatic strain $\varepsilon_{\text{hydro}}$ in a NW with 25% In in the shell. $\varepsilon_{\text{hydro}}$ is dominated by the principal component along the NW longitudinal axis, as reported elsewhere [40]. Shear strain components are one order of magnitude smaller than $\varepsilon_{\text{hydro}}$, in
agreement with previous reports [40]. As expected from the core-shell lattice mismatch, $\epsilon_{\text{hydro}}$ is on average negative in the shell (compressive) and positive in the core (tensile). We find about 2.2% of tensile strain [position “A” in Fig. 3(b)], softly modulated following a sixfold symmetry. The shell is subject to a maximum compression of $-0.35\%$ in the middle of the NW facet [position “C” of Fig. 3(b)]. The shell strain intensity is modulated by a sixfold symmetry in a more pronounced way. The outer corners are relaxed [minimum strain of 0% at position “B” in Fig. 3(b)]. Importantly, the regions of smaller strain in the shell expand around the NW sidewall ridges, becoming larger the farther they are from the core. The experimental (PL) [41] and simulated strain values are plotted in Fig. 3(c): the simulated compressive strain in the shell (position C) and the experimental and simulated tensile strain in the core increase linearly with the In fraction. The theoretical and experimental values agree within 15%.

A growth model that explains the nonrandom distribution of In and Al in the shell is presented. The expected behavior of Ga and In adatoms at the NW facets is sketched in Fig. 3(d). We start by addressing the Al distribution. Al tends to incorporate more efficiently at the vertices of the hexagonal cross section ([112] facets) [15,18,19]. This is the consequence of a larger sticking coefficient on those facets [42,43]. For a similar reason, it has been shown that the incorporation is higher on B- than A-polar facets [18]. This results in a threefold symmetry of the Al distribution.

We now turn to the incorporation of In. According to our simulations, there is a compressive strain in the middle of the facets (position C in the drawings). Since AlGaNAs exhibits a larger lattice constant than AlGaAs, we expect In to be favorably incorporated in the relaxed corners. In adatoms exhibit high mobilities, allowing diffusion to occur at the scale of the NW facets. Strain relaxation should support an In flux towards the more relaxed [112] corner nanofacets, while the [110] facets exhibit a significantly lower In fraction in the center [arrows leaving this position in Fig. 3(d)]. Qualitatively, there is good match between the wedge-shaped In segregation and the simulated strain distributions that show increasing strain relaxation towards the outer NW shell. This correlation suggests that strain has an important role in shaping the In segregation. It is also known that moving steps in step-flow or 2D-island growth may drive the atom segregation [44]. Here, it is difficult to assess how atomic steps affect the In diffusion without performing in-situ experiments. However, the steps may enhance, but definitely do not significantly prevent, the strain-driven segregation of In at the NW ridges. In highly lattice-mismatched GaAs-InGaAs core-shell NWs [5], the coherent shell deposition on the [110] sidewalls occurs simultaneously with the strain-driven migration of In-rich material towards protrusions at the NW ridges. Therefore nanoscale strain gradients are the major force driving the In segregation at the ridges of lattice-mismatched NW heterostructures.

We also observe that In is poorly incorporated at the highly Al-rich stripes. We attribute this to the low miscibility of In with Al-containing alloys [45]. At the [112] nanofacets, In competes with Al for the available binding sites. This is particularly relevant for the B-polar nanofacets because Al tends to accumulate more in these positions. Throughout the shell, the In segregation strongly correlates with polarity: regions with a markedly higher In fraction coincide with the B-polar vertices. We note that the polarity-driven incorporation of In has not been demonstrated in the past. The competition between In and Al to stick on the [112] facets strongly favors Al on the B-polar surfaces; this mechanism may generate a net flux of In atoms towards the surroundings of the [112] B Al-rich planes. The literature on GaAs-AlGAs core-shell NWs also reports that the [112] facets are systemati-

Conclusion. In conclusion, we have demonstrated a significant (300 meV) redshift of the emission energy of single-photon emitters embedded in the shell of core-shell GaAs-AlGaNAs NWs. The redshifting mechanism is based on the In incorporation in the AlGAs shell alloy to form a quaternary semiconductor of lower band gap. The spatial distribution of different species in the shell is determined by several high-resolution techniques of compositional analysis. In addition to the well-known Al and Ga segregation, we find evidence of novel wedge-shaped In-rich segregation. We explain the distribution of the different species in the quaternary alloy with their diffusion on the NW sidewalls driven by crystal phase, polarity, and strain. Finite-element simulations provide insight in the role of strain to drive the segregation of In. These findings advance the understanding of the segregation phenomena in quaternary alloys, as required to take full advantage of the additional degrees of freedom that they offer.

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